

Nuclear level densities and γ -ray strength functions of $^{180,181,182}\text{Ta}$ and nucleosynthesis of ^{180}Ta

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Introduction

Physics Motivation for ^{180}Ta

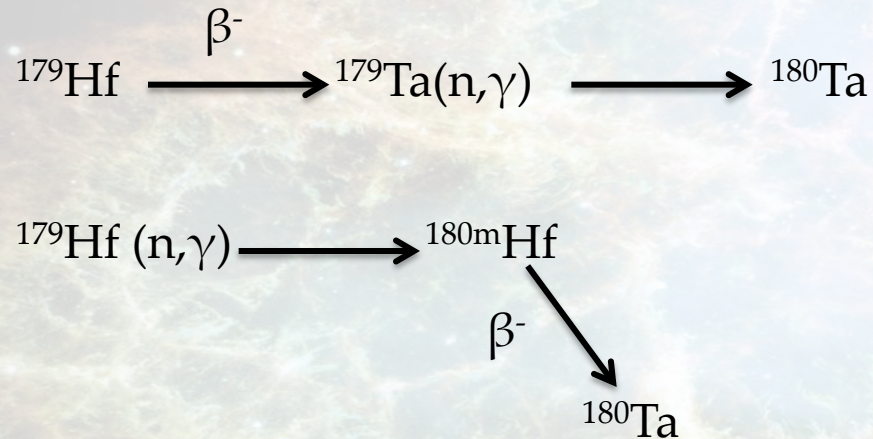
- Almost all $A > 110$ p -nuclei are thought to be produced by the photodisintegration of s - and r - process pre-existing nuclei
- However, the observed low abundant (0.012%) ^{180}Ta remain an exception
- Several processes are considered
- ^{180}Ta could be explained with p -process [1]

[1] M. Arnould and S. Goriely, Phys. Rep. 384, 1 (2003).

Introduction

Physics Motivation for ^{180}Ta

- Several processes are considered
- ^{180}Ta could be explained with p-process [1]
- The s-process mostly via branching in ^{179}Hf [2]



[1] [1] M. Arnould and S. Goriely, Phys. Rep. 384, 1 (2003).

[2] M. Loewe et al. Nucl.Phys. A719 (2003) 275c

Introduction

Physics Motivation for ^{180}Ta

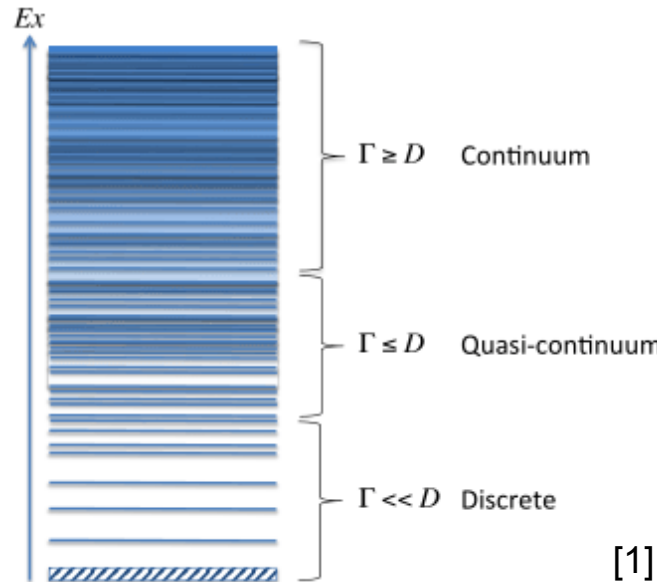
- $^{180}\text{Hf}(\nu_e, e)^{180}\text{Ta}$, $^{181}\text{Ta}(\nu, \nu'n)^{180}\text{Ta}$ and p -process contribute 50, 25 and 25% respectively [3]
- $^{180}\text{Hf}(\nu_e, e)^{180}\text{Ta}$ 50% contribution recently supported by A. Byelikov et al. [4]
- This is unsettled, hence provoking debates
- Uncertainties in reaction rates for ^{180}Ta rest in nuclear properties
- Therefore synthesis of ^{180}Ta needs to be reinvestigated

[3] A. Heger et al. Phys.Lett. B 606, 258 (2005)

[4] A. Byelikov et al. PRL 98, 082501 (2007)

Objectives

- Nuclear level density (NLD) and γ -ray strength function (γ SF) below S_n in $^{180,181,182}\text{Ta}$ isotopes (Oslo Method)



- Astrophysical Maxwellian-averaged (n, γ) cross sections (TALYS)
- Investigate production mechanism of ^{180}Ta

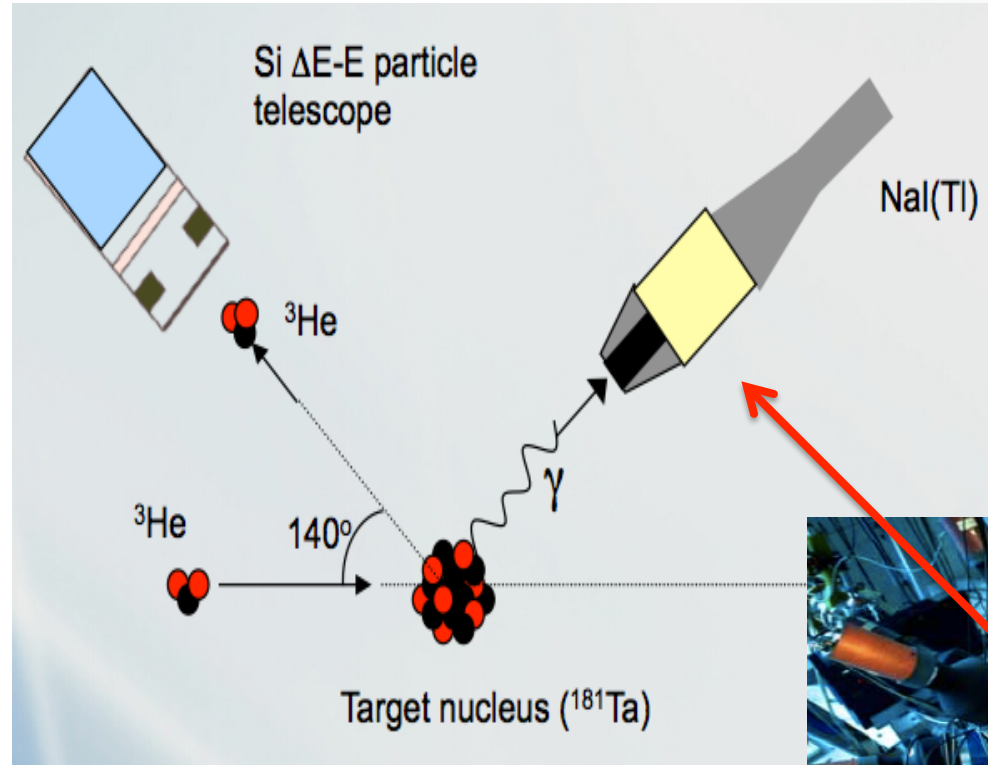
[1] Magne Guttormsen et al.: Experimental level densities of atomic nuclei

Experimental Details

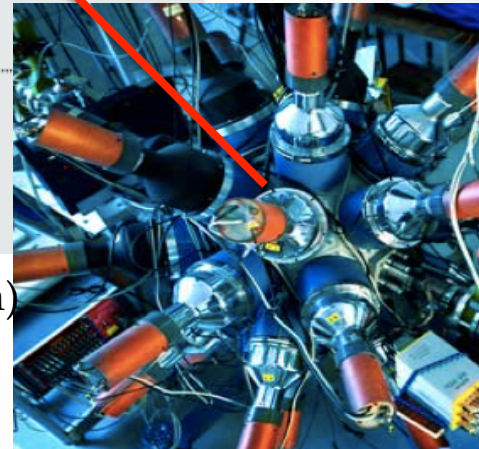
Experimental setup at the Oslo cyclotron laboratory



- 0.8 mg/cm² thick ¹⁸¹Ta target
- ¹⁸¹Ta(³He,X)^{180,181}Ta, 34 MeV
- ¹⁸¹Ta(d,X)^{180,181}Ta, 15 MeV
- ¹⁸¹Ta(d,X)^{181,182}Ta, 12.5 MeV



- CACTUS Array: 26 collimated 5"×5" NaI(Tl) (~22 cm)
- 14.1% eff. at E_γ= 1332 keV



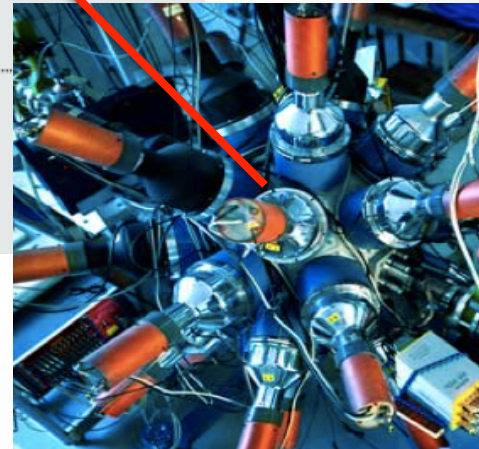
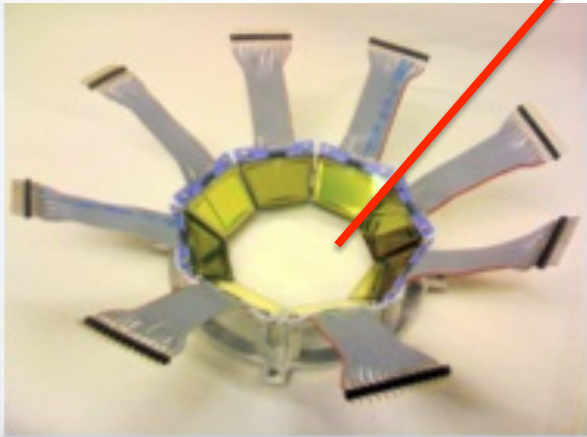
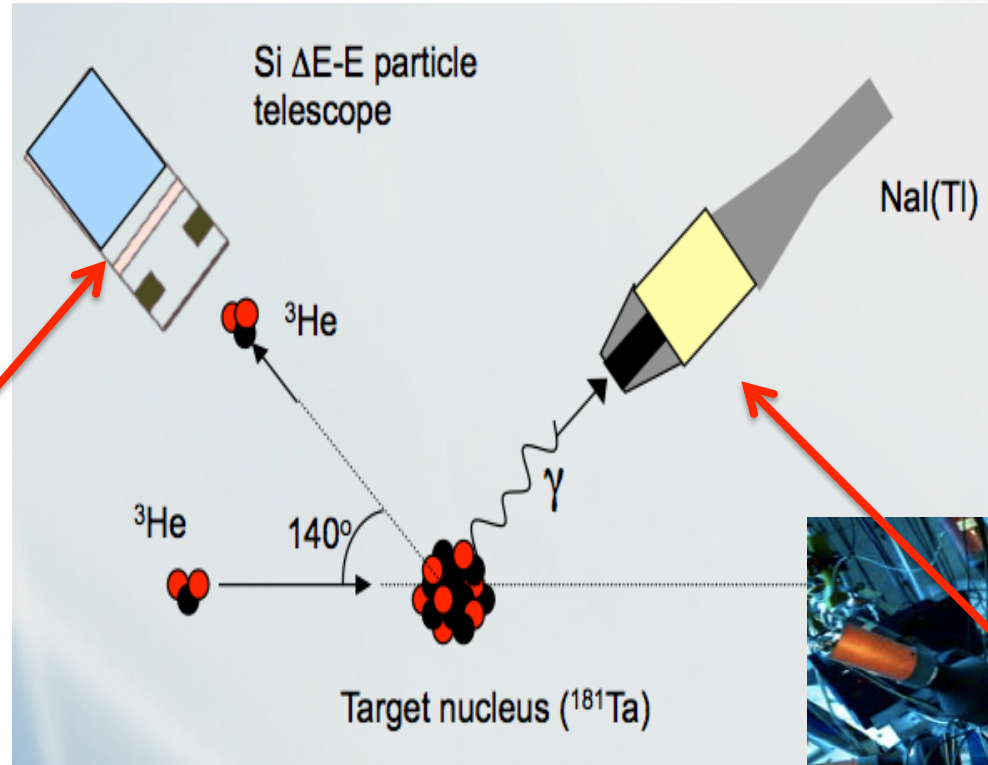
Experimental Details

Experimental setup at the Oslo cyclotron laboratory



- 0.8 mg/cm² thick ¹⁸¹Ta natural target

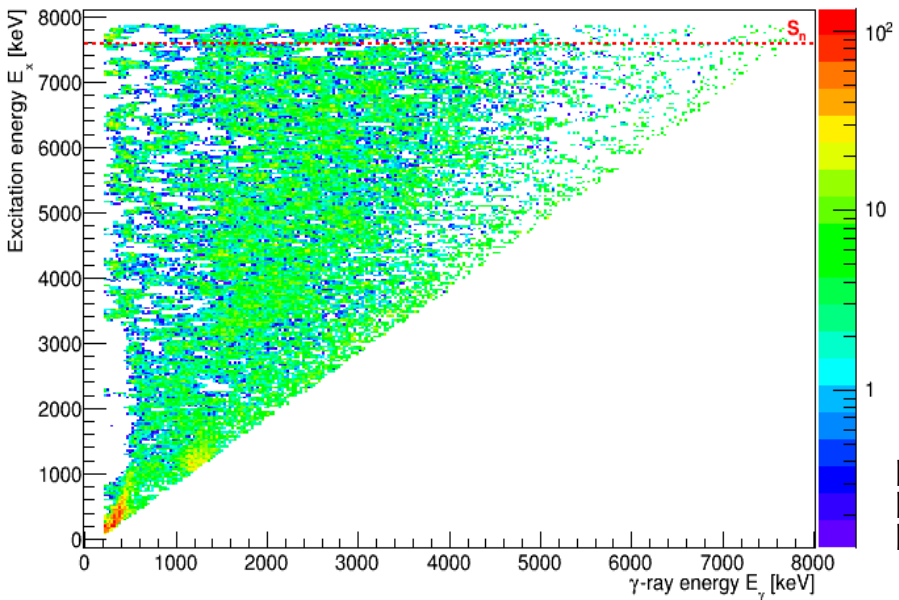
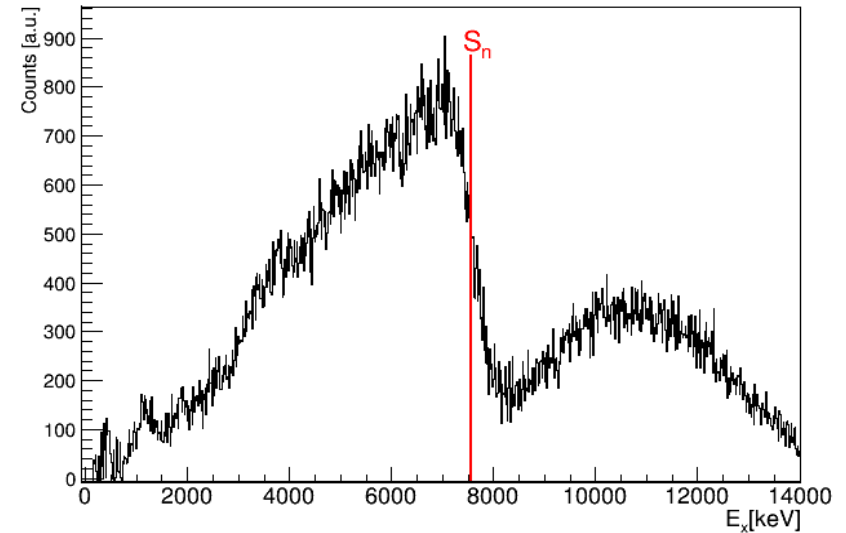
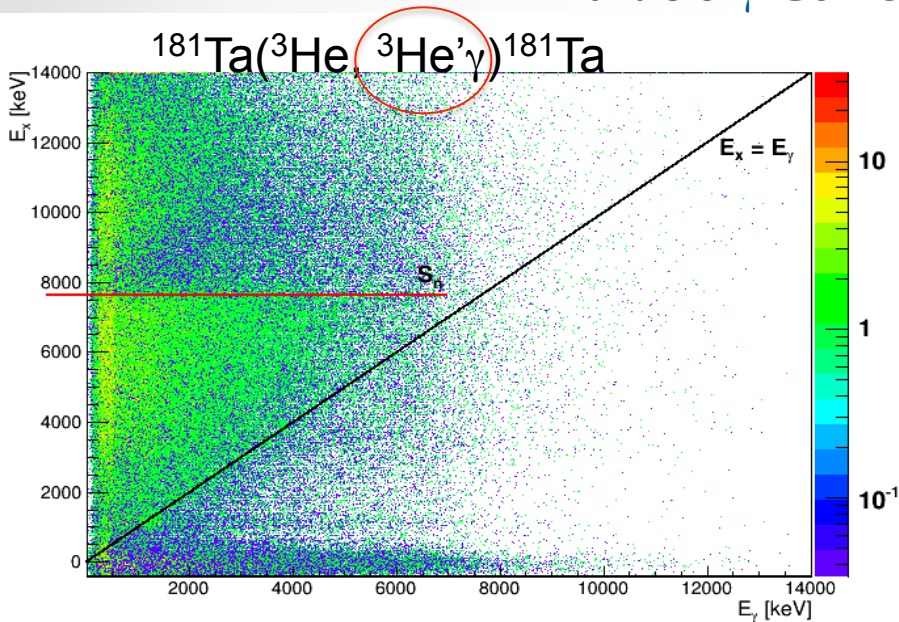
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- SiRi Array, 64 ΔE -E Si particle telescopes
- ΔE , E and Al foil thicknesses (130 μm , 1550 μm and 10.5 μm)
- $\theta = 126^\circ$ to 140° , ~ 5 cm

Data Analysis

Particle- γ Coincidence Matrices



The Oslo Method

1. Unfolding γ -ray continuum spectra[1]
> Unfolding iterative procedure
2. Extraction of primary γ -rays [2]
> first-generation method
3. Simultaneous extraction of level density and strength function[3]
> Assumes a compound state

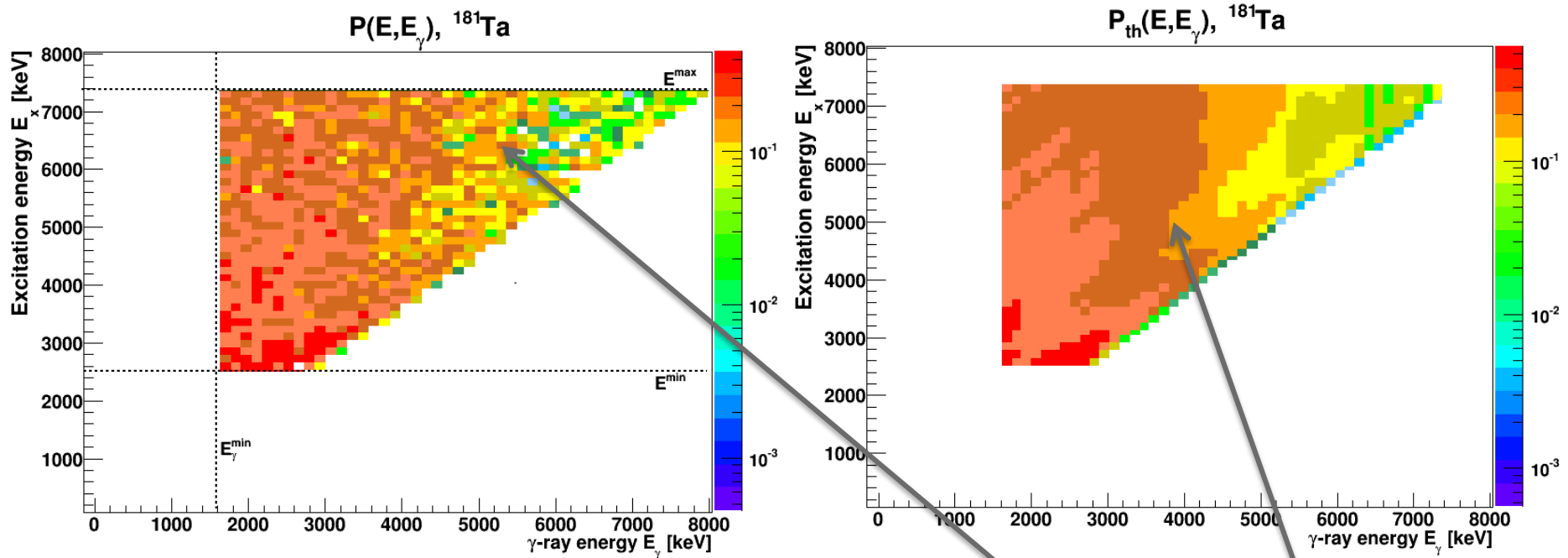
[1] M. Guttormsen et al., NIM Phys. Res. A 374, 371 (1996)

[2] M. Guttormsen et al., NIM Phys. Res. A 255, 518 (1987)

[3] A. Schiller et al., NIM Phys. Res. A 447, 498 (2000)

Data Analysis

Experimental vs. Fitted FG Matrices



$$P(E, E_\gamma) \propto \rho(E_f) T(E_\gamma).$$

- Assumptions: Brink Hypothesis
- The $\rho(E_f)$ and $T(E_\gamma)$ are extracted by fitting to $P(E_x, E_\gamma)$ and performing χ^2 minimum.

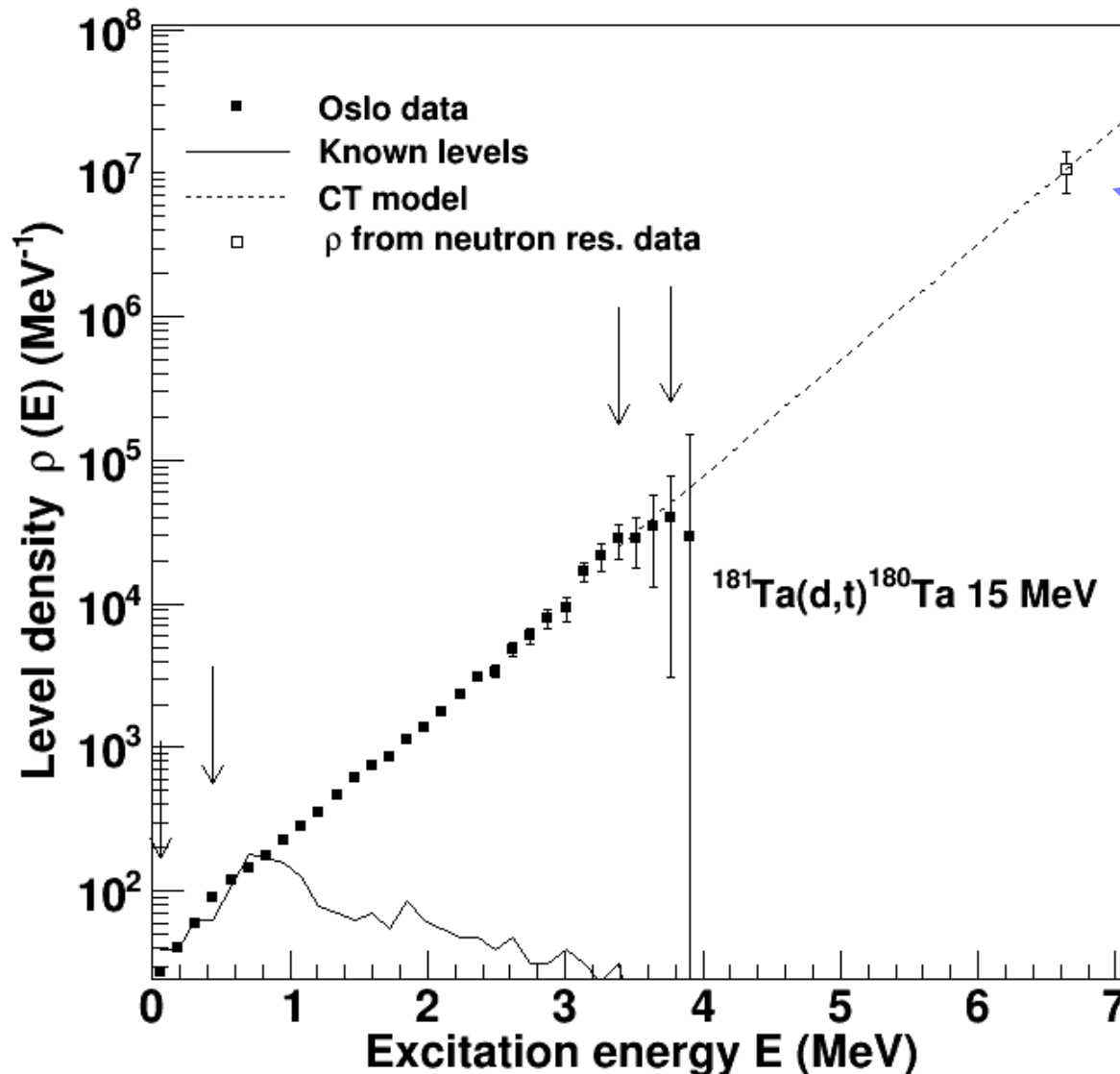
The Oslo Method

- Normalization of (γ transmission coefficient) $T(E_\gamma)$ and $\rho(E_f)$
- Calculation of γ SF from $T(E_\gamma)$

^{180}Ta Results

Normalization of the NLD

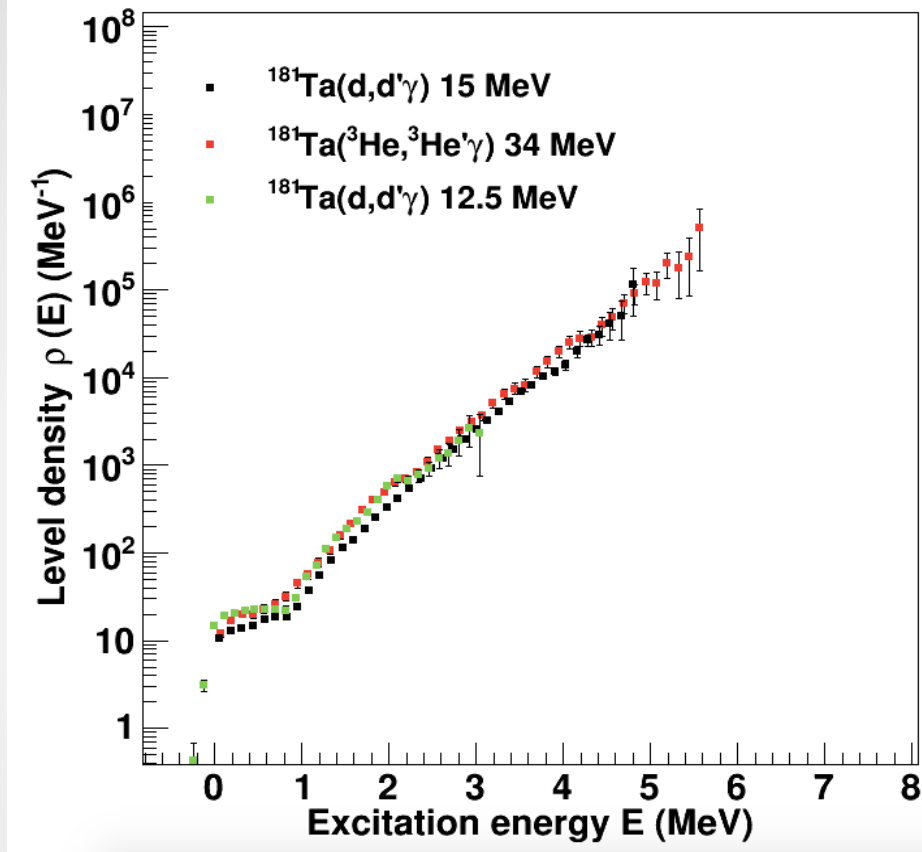
$$\rho(S_n) = \frac{2\sigma^2}{D_0(J_T + 1)e^{[-(J_T+1)^2/2\sigma^2]} + e^{(-J_T^2/2\sigma^2)}J_T}$$



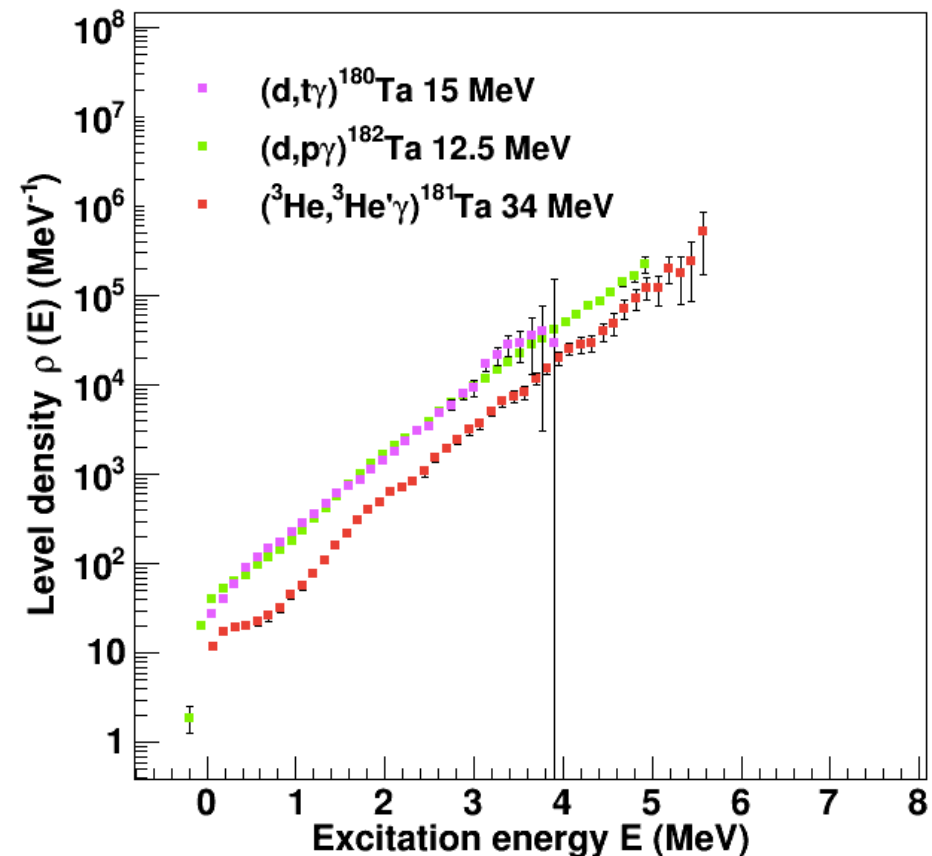
No neutron resonance data available (^{180}Ta)

Nuclear Level Densities

Different reactions yield similar results

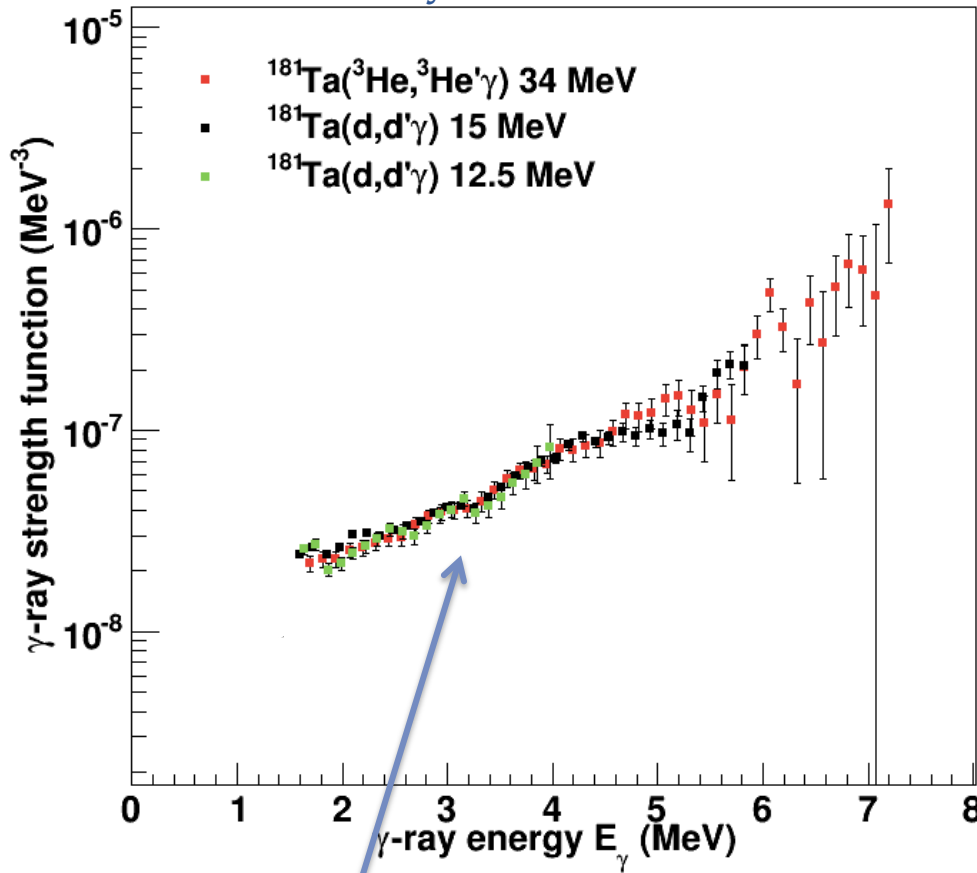


NLDs of neighbouring isotopes



γ -ray Strength Functions

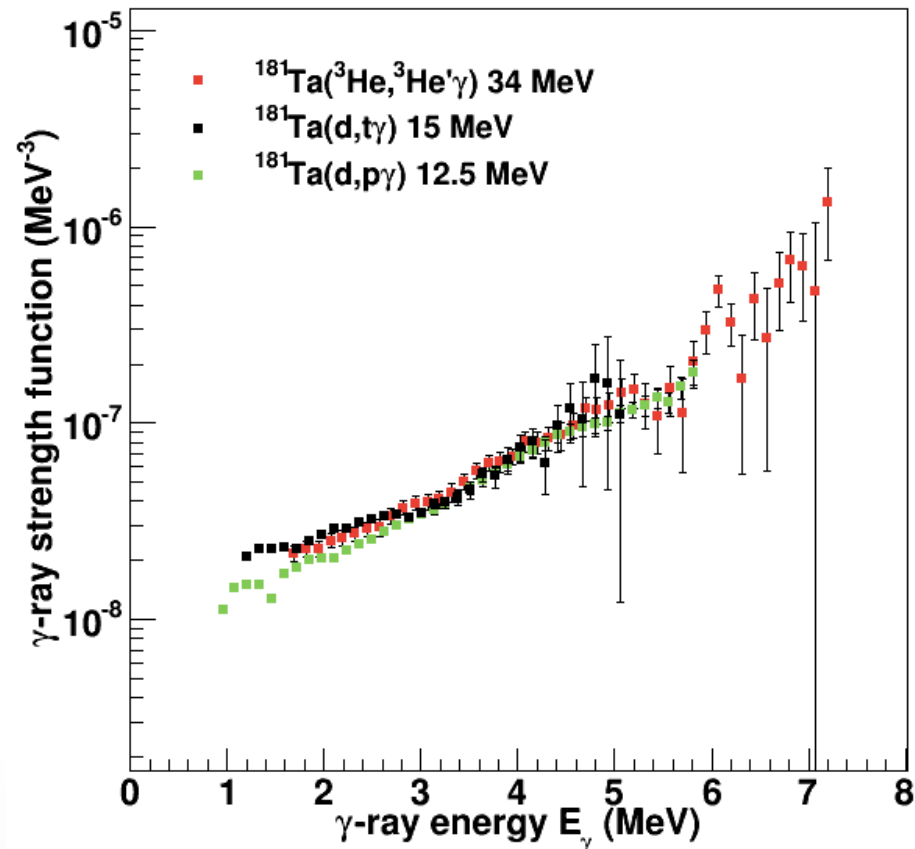
Different reactions yield similar results



C.P Brits (Poster: Search for scissors resonance in $^{181,182}\text{Ta}$) 15/09/2016

$$f(E_\gamma) = \frac{1}{2\pi E_\gamma^3} B\mathcal{T}(E_\gamma) \quad [1]$$

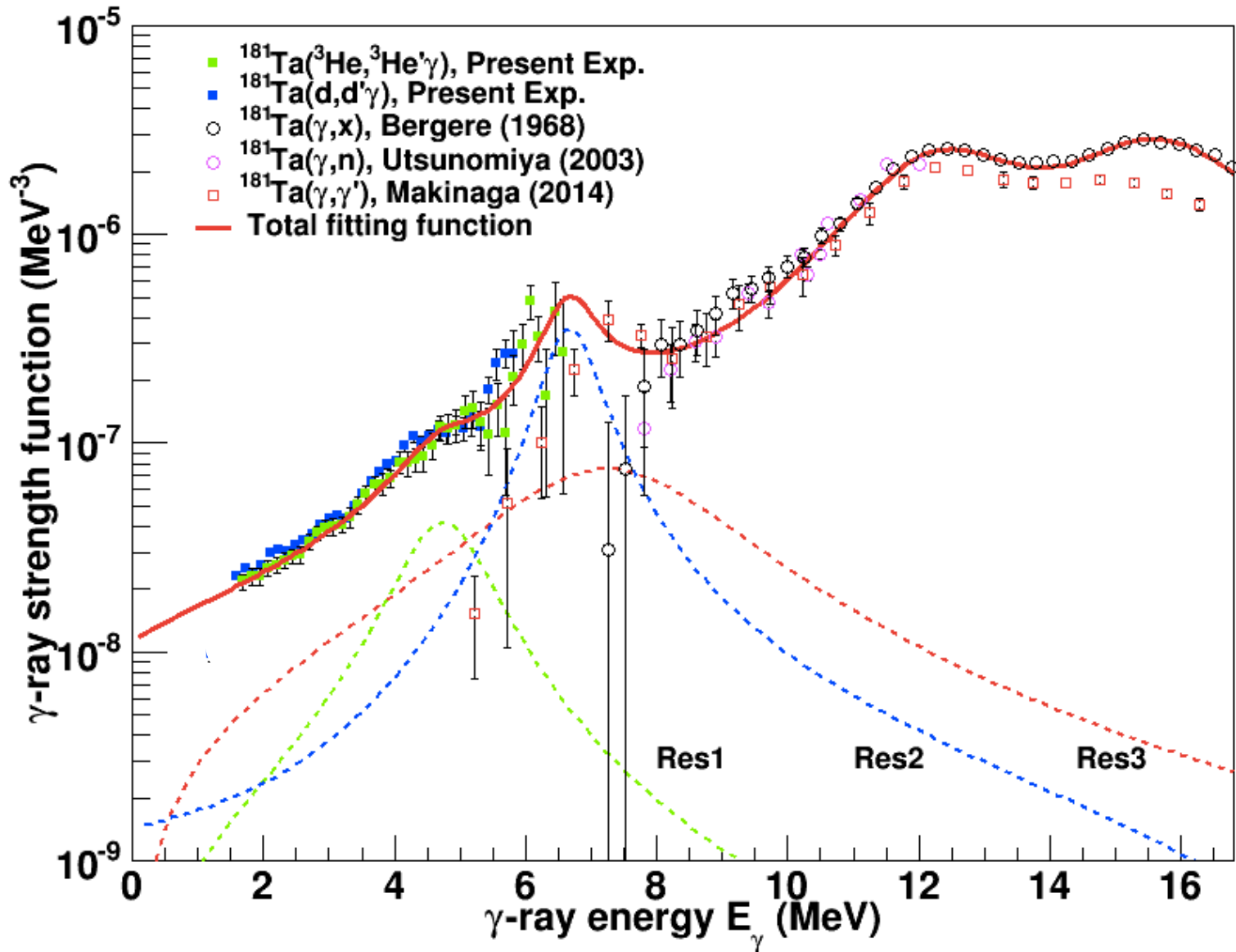
γ SFs of neighbouring isotopes



^{181}Ta Results

Comparison of $^{181}\text{Ta}(^3\text{He}, ^3\text{He}'\gamma)$ and $^{181}\text{Ta}(d, d'\gamma)$ γ SF with photoabsorption cross-section data

$$f(E_\gamma) = \frac{1}{3\pi\hbar^2 c^2} \left[\frac{\sigma(E_\gamma)}{E_\gamma} \right]$$



^{182}Ta Preliminary Results

Neutron capture Cross Section Calculations

Upper error bands

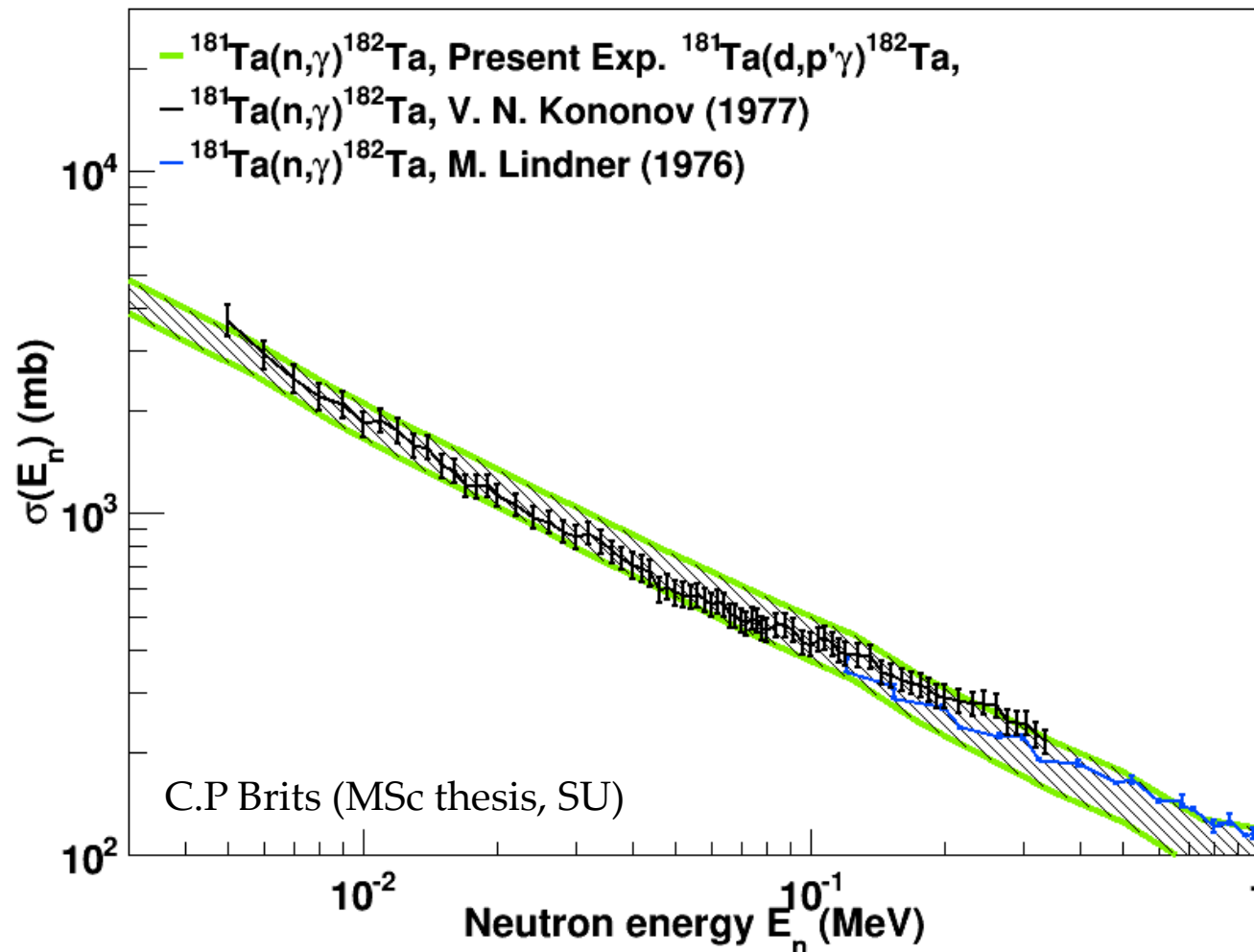
$$D_0 = D_0 - \delta D_0$$

$$\langle \Gamma(S_n) \rangle = \langle \Gamma(S_n) \rangle + \langle \delta \Gamma(S_n) \rangle$$

Lower error bands

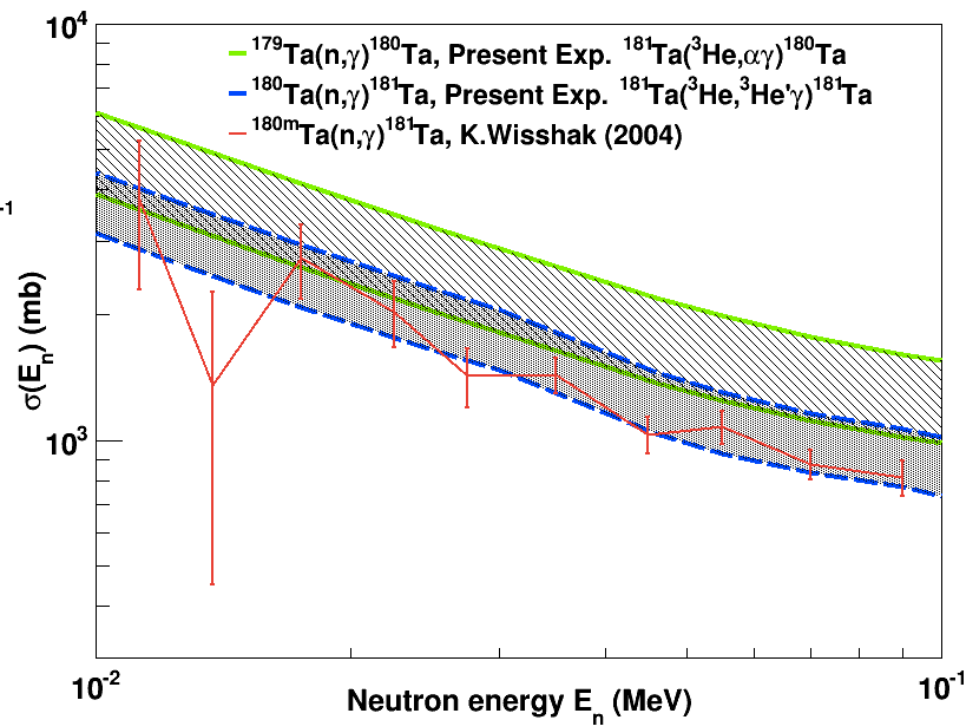
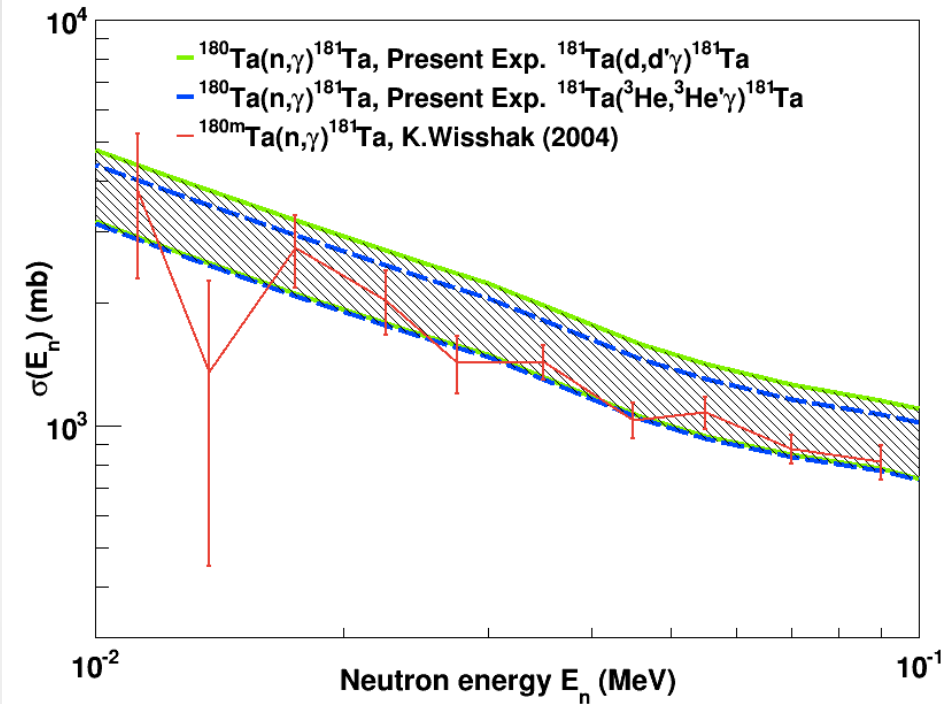
$$D_0 = D_0 + \delta D_0$$

$$\langle \Gamma(S_n) \rangle = \langle \Gamma(S_n) \rangle - \langle \delta \Gamma(S_n) \rangle$$



Preliminary Results

Neutron capture Cross Section Calculations



180,181,182Ta Preliminary Results

Maxwellian-Average Cross Sections (MACS)

Reaction	$\langle E \rangle$ (keV)	MACS (mb)	Rec. MACS (mb)
$^{179}\text{Ta}(n,\gamma)$	30	2469 ± 508	1334 ± 422 [1]
	215	828 ± 288	
$^{180}\text{Ta}(n,\gamma)$	30	2051 ± 148	1465 ± 100 [2,3]
	215	589 ± 53	
$^{181}\text{Ta}(n,\gamma)$	30	950 ± 136	766 ± 15 [4]
	215	259 ± 46	

[1] Z. Y. Bao et al. , Atomic Data and Nuclear Data Tables 76 , 70 (2000).

[2] K. Wisshak et al. , Phys. Rev. C 69 , 055801 (2004).

[3] K. Wisshak et al. , Phys. Rev. Lett. 87 , 251102 (2001).

[4] K. Wisshak, F. Voss, F. Käppeler, and G. Reffo, Phys. Rev. C 42, 1731 (1990).

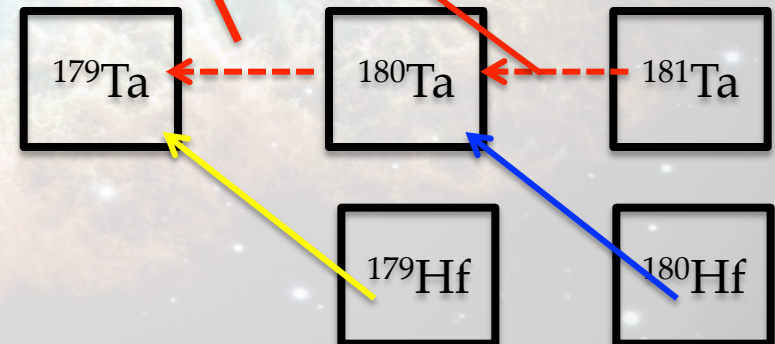
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Assumption: Detailed balance

s-process??



Summary And Future Plans

- Experiments successfully performed at the Oslo cyclotron laboratory
 - $^{181}\text{Ta}(3\text{He},\text{X})^{180,181}\text{Ta}$, 34 MeV
 - $^{181}\text{Ta}(\text{d},\text{X})^{180,181}\text{Ta}$, 15 MeV
 - $^{181}\text{Ta}(\text{d},\text{X})^{181,182}\text{Ta}$, 12.5 MeV
- γSF and NLD of $^{180,181,182}\text{Ta}$ extracted with the Oslo Method
- First time, measurements of NLD and γSF below Sn in ^{180}Ta and below 5 MeV in ^{181}Ta experimentally.
- The experimental γSF and NLD used to investigate (n,γ) cross sections
- The newly deduced $^{179,180}\text{Ta}(\text{n},\gamma)$ cross sections are to be utilized in astrophysical network calculations to investigate and constrain the galactic production mechanism of ^{180}Ta
- C.P Brits (Poster: Search for scissors resonance in $^{181,182}\text{Ta}$) 15/09/2016

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Thank You For Your Attention !!!

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